# Reconstructing balloon-observed gravity wave momentum fluxes using machine learning and input from ERA5

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# Key Points:

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12	•	Eight superpressure balloons from the Strateole 2 mission provide observations
13		for accurate gravity wave momentum flux estimation
14	•	Three machine learning methods are employed to probe the relationship between
15		the gravity wave momentum fluxes and ERA5's large-scale flows
16	•	The most informative large-scale inputs are provided, along with a discussion of
17		the successes and challenges of machine learning methods

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## 18 Abstract

Global atmospheric models rely on parametrizations to capture the effects of gravity waves 19 on middle atmosphere circulation. The momentum fluxes associated with these waves, 20 as they propagate upwards from the troposphere, represent a crucial yet insufficiently 21 constrained component. This study employs machine learning (ML) techniques to probe 22 the relationship between large-scale flows and small-scale gravity waves within the trop-23 ical lower stratosphere. The measurements collected by eight superpressure balloons from 24 the Strateole 2 campaign, comprising a cumulative observation period of 680 days, pro-25 vide valuable reference estimates of the gravity wave momentum fluxes. Multiple explana-26 tory variables, including total precipitation, wind, and temperature, were interpolated 27 from the ERA5 reanalysis at each balloon's location. Three tree-based ensemble ML meth-28 ods are trained using data from seven balloons, and subsequently applied to estimate the 29 reference gravity wave momentum flux for the remaining balloon. The numerical results 30 show that parts of the gravity wave signal are successfully reconstructed, with correla-31 tions typically around 0.54 and exceeding 0.7 for certain balloons. In addition, the per-32 formances of the ML models exhibit greater sensitivity to the selection of training data 33 than to the ML method. The most informative ERA5 inputs generally include total pre-34 cipitation and wind variables near the balloons' level. However, two distinct methods 35 achieving similar accuracy, may favor different flow variables. This study also discusses 36 potential limitations of ML methods, such as the intermittent nature of gravity wave mo-37 mentum fluxes and data scarcity, providing insights into the challenges and opportuni-38 ties for advancing our understanding of these atmospheric phenomena. 30

#### <sup>40</sup> Plain Language Summary

Part of the atmosphere's large-scale circulation results from motions that are not 41 resolved, or partly resolved, by weather or climate models. These include internal grav-42 ity waves (GWs), with horizontal scales from a few to hundreds of kilometers. The main 43 sources occur in the troposphere, such as flow over mountains and cloud development. 44 Their three-dimensional propagation induces major aggregated impacts in the strato-45 sphere and mesosphere, forcing key aspects of the circulation. This forcing is accounted 46 for in climate models by 'parameterizations', that mimics the effect of the unresolved waves 47 based on the large-scale, resolved flow. These parameterizations necessarily retain crude 48 approximations and introduce significant uncertainty in the models. For GWs, sources 49 are a major uncertainty. This study makes use of the high-altitude balloon campaign Stra-50 teole 2 (Oct. 2019-Feb. 2020). Eight balloons circled Earth at heights around 18 to 20 51 km, providing unique observations of the GWs. These are used as targets for machine 52 learning (ML) methods that take as inputs the information from outputs of a numer-53 ical weather prediction model describing the large-scale flow. The successes and difficul-54 ties of ML provide insights which can guide improvements of parameterizations, such as 55 the most informative large-scale variables for estimating the unresolved waves. 56

# 57 1 Introduction

Climate models and Numerical Weather Prediction models resolve a widening range 58 of atmospheric processes as computing power increases, enabling finer spatial resolution. 59 Subgrid-scale processes persist nonetheless, and efforts to improve and constrain them 60 better are essential. Internal gravity waves constitute one of these subgrid-scale processes, 61 with important implications for the circulation and variability of the middle atmosphere 62 (Fritts & Alexander, 2003). Motivations for improved modeling of the stratosphere in-63 cludes climate (e.g. Solomon et al. (2010); Kremser et al. (2016)) but also predictabil-64 ity on shorter time scales (F. Vitart and A.W. Robertson, 2018; Butchart, 2022). 65

Gravity waves occur on scales ranging from a few to several hundreds of kilometers. An important effect stems from their vertical propagation: gravity waves are re-

sponsible for vertical transfers of momentum from lower layers (troposphere: denser and 68 with more gravity wave sources) to upper layers (stratosphere and beyond), where they 69 constitute an essential driver of the overall circulation (Fritts & Alexander, 2003). A sig-70 nificant part of the spectrum of gravity waves has been and remains unresolved in global 71 models, requiring these effects to be represented by parameterizations (Kim et al., 2003). 72 Models display sensitivity to these, calling for coordinated efforts to better constrain these 73 parameterizations from both observations and high-resolution modeling (Alexander et 74 al., 2010). 75

76 A global comparison of observed, resolved and parameterized gravity wave momentum fluxes (GWMFs) was carried out by Geller et al. (2013), highlighting significant dis-77 crepancies. Although GWs parameterizations are now used routinely in climate mod-78 els, their validation against in situ obestivations remain a challenge. There exist global 79 observations derived from satellite observations (e.g. Ern et al. (2018)), but there are 80 limitations on the wavelengths that can be observed, and significant assumptions are needed 81 to indirectly deduce important quantities like the momentum fluxes from temperature 82 fluctuations, using polarization relations (Alexander et al., 2010; Ern et al., 2014). For 83 these reasons superpressure balloons have been highlighted as a valuable and accurate 84 source of information on gravity wave momentum fluxes (Geller et al., 2013). A down-85 side of superpressure balloon observations is their very sparse sampling of the lower strato-86 sphere: despite a broad coverage of the Southern Ocean (Jewtoukoff et al., 2015) and 87 of the equatorial belt (Corcos et al., 2021), each balloon flight provides only local infor-88 mation: one time series along its trajectory. 89

There are fundamental difficulties in validating parameterizations of gravity waves: 90 the purpose of a parameterization is to provide the forcing to the large-scale which is miss-91 ing because of unresolved processes. Ideally, one would wish to know what this forcing 92 should be and validate this outcome of parameterizations. Unfortunately, this forcing 93 can not be directly observed. Validating parameterizations by the realism of the clima-94 tology and variability of the atmospheric circulation in global models constitutes a first 95 step, but is not a severe test and allows for compensating errors between parameterized 96 processes (Plougonven et al., 2020). More stringent tests involve comparisons to obser-97 vations (de la Camara et al., 2014; Trinh et al., 2016). Recently, direct comparisons be-98 tween observed and parameterized gravity waves have been carried out on the scale of 99 daily variations rather than at the level of general statistical characteristics (Lott et al., 100 2023). The large-scale environment was described using the ERA5 reanalyses (Hersbach 101 et al., 2020), providing the background fields necessary to emulate the parameterization 102 of convectively generated waves of Lott & Guez (2013). The comparison was quite en-103 couraging, with the gravity wave momentum fluxes having the right order of magnitude, 104 and an appropriate intermittency. 105

An essential aspect, and fundamental issue, to keep in mind when comparing ob-106 served and modeled gravity wave momentum fluxes is their strong intermittency: in time-107 series of GWMF, one commonly finds short, intense peaks corresponding to a strong grav-108 ity wave event, surrounded by considerably weaker values. This has been highlighted in 109 the long 'tail' of the Probability Density Function (PDF) of the GWMF (Alexander et 110 al., 2010; Hertzog et al., 2012), and quantified in simulations and observations (Plougonven 111 et al., 2013; Wright et al., 2013). This intermittency further contributes to making the 112 parameterization of gravity waves a challenging task. 113

For the improvement of parameterizations in general (not only those of gravity waves), machine learning methods provide an array of possibilities. These have been explored in different directions:

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- Machine learning can enable the emulation of parameterizations, leading to significant computational time savings (Chantry et al., 2021).

- Machine-learning can help to capture the relationship between large-scale fields and the unresolved process, as illustrated in the case of convection by Gentine et al. (2018). For exploration, the dataset used as the truth came from a higher-resolution simulation, not from observations; obtaining observationally based knowledge of the effects to be parameterized remains a major challenge.
  Machine learning can be used to explore the relationship between the large-scale flow and the resulting small-scale waves, as has been done for orographic waves
- 125 flow and the resulting small-scale waves, as has been done for orographic waves 126 over Northern Japan (Matsuoka et al., 2020). Again, both the target and the in-127 puts are modelled fields, but at different resolutions.
- As a preliminary to a data-driven parameterization that would have learned from observations, a machine-learning-based emulator of a parameterization for gravity waves has been used in a climate model, including under climate change conditions (Espinosa et al., 2022).

The purpose and scope of the present study is to probe the relationship between 132 the large-scale flow and gravity waves in the Tropics, using machine learning approaches 133 to address fundamental issues: what fraction of the GWMF can be determined from knowl-134 edge of the large-scale, and what fraction remains as *stochastic*? Which large-scale vari-135 ables are most informative, and do they match with our common understanding of un-136 derlying gravity wave parameterizations? The present study belongs to the third cat-137 egory outlined above for the uses of machine learning (the purpose is *not* to produce a 138 new parameterization, nor to emulate an existing one). With similar goals, Amiramjadi 139 et al. (2023) used Machine Learning methods to probe the relationship between the large-140 scale flow and gravity waves, for non-orographic waves in the mid-latitudes and using 141 waves resolved in a reanalysis as a target. In contrast, the present study aims at observed 142 momentum fluxes in the Tropics, where the Strateole 2 campaigns provide a wealth of 143 new observations (Haase et al., 2018; Corcos et al., 2021). 144

The paper is organized as follows: Section 2 provides an overview of the data and ML algorithms used in this study. Section 3 presents the performances of ML methods in reconstructing the reference GMWFs. Section 4 discusses the factors that influence the performances and addresses the limitations of ML methods. Finally, Section 5 concludes the study with key takeaways and future directions.

## <sup>150</sup> 2 Data and methodology

#### 2.1 Data

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We use in situ observations collected from eight constant-level balloon flights (altitude between 18.5 and 20km) during the Strateole-2 mission from November 2019 to February 2020 (Corcos et al., 2021).

As in Corcos et al. (2021), momentum fluxes (MFs) were computed from raw bal-155 loon measurements following the procedure described in Vincent and Hertzog (2014). Es-156 sentially, the pressure and horizontal wind time series are first projected in the time-frequency 157 domain thanks to a continuous wavelet transform (Torrence and Compo, 1998). The pres-158 sure observations inform on the vertical displacements of the balloon, which are related 159 to those of air parcels, assuming that the balloon behaves as a perfect isopycnic tracer. 160 The time-frequency MF decomposition is then derived from the wavelet cross-spectrum 161 of the horizontal winds and air-parcel vertical displacements. Segments polluted by non-162 geophysical artifacts (e.g. depressurization events) are discarded. 163

For our analysis, and following Corcos et al. (2021), we considered gravity wave MFs integrated over two frequency bands: a high-frequency (HF) band (i.e. short periods, ranging from 15 minutes to 1 hour) and wide-frequency (WF) band (i.e., long periods, ranging from 15 minutes to 1 day). Additionally, we also differentiate between eastward-

propagating waves that yield positive MF in the zonal direction (eastward) and westward-168 propagating waves that produce negative MF (westward). We use these MFs as a ref-169 erence for the true target MFs. Then, we pair them with large-scale flow input informa-170 tion from ERA5. These input variables are sampled at 5 by 5 horizontal grid, with each 171 grid cell having a resolution of  $1^{\circ} \times 1^{\circ}$ . The grid spans across 67 vertical levels, consist-172 ing of all odd levels within the 137-level model. It shifts along the balloon trajectories, 173 with the center being the nearest point to the balloon position. The inputs and the tar-174 gets are interpolated and averaged into 1-hour time resolution. The three ML models 175 are trained using 3-hour time averaging data, and their performance will be evaluated 176 based on daily averaging time resolution, as presented in Lott et al. (2023). 177

#### 178 2.2 Methodology

In this study, three tree-based ensemble ML methods are considered: random forest (RF) introduced in Breiman (2001), extremely randomized trees also known as extratrees (ET) by Geurts et al. (2006), and Adaptive Boosting or Adaboost regressors by Freund & Schapire (1997). These algorithms construct multiple decision trees, and the final prediction is determined by aggregating the individual decision tree predictions.

It should be noted that other methods, such as deep neural networks, as well as other types of networks including convolutional and recurrent neural networks, have also been implemented. However, the performances of these methods are not comparable to the presented tree-based algorithms, as these models typically require a large number of observations to achieve comparable results. The limitations and concerns regarding the models, the large-scale input variables, the target observations, and the nature of the relation between the large-scale and small-scale flows will be discussed later in Section 4.3.

#### 2.2.1 Decision tree

The decision tree algorithm (Breiman et al., 1984) is the foundational building block of the primary ML methods used for our predictions. They are widely used for nonlinear prediction problems due to their efficiency and interpretability. To construct a decision tree, the training data is recursively partitioned into small hyperrectangular regions of the forms  $R_1 = \{X \leq \alpha\}$  and  $R_2 = \{X > \alpha\}$  for some ERA5 input variable X (wind velocity or precipitation, for instance) and threshold  $\alpha$ . At each step, we recursively split the input space into hyperrectangular regions that are as pure as possible. Purity refers to the homogeneity of the target y (GWMF) within each region, and Total Within Sum of Squares (TWSS) is utilized as the impurity measure in this study. Specifically, a split is performed at any input variable X at threshold  $\alpha$  if it minimizes the following TWSS criterion:

$$\sum_{y \text{ of } R_1} (y - \mu_1)^2 + \sum_{y \text{ of } R_2} (y - \mu_2)^2$$

#### 192 where

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•  $R_1$  and  $R_2$  are the left and right regions of the split

•  $\mu_1$  and  $\mu_2$  are the average target within region  $R_1$  and  $R_2$  respectively.

Any new observation must belong to one of these regions, and its prediction is determined by averaging the target values of all the neighboring observations within that block. Constructing an optimal tree is generally challenging, and the tree's structure, such as its depth and the minimum size of regions allowed to split, are hyperparameters that need to be optimized. Figure 1 below provides an example of a simple decision tree trained on 100 observations of precipitation and zonal wind velocity to predict absolute GWMF.



**Figure 1.** An example of a simple decision tree built using precipitation and wind velocity to predict absolute GWMF. The left side is the partition cell representation of the tree on the right side. The data points are colored according to the value of its target GWMF.

#### 202 2.2.2 Random forest

Random forest (RF) is a powerful ensemble learning method that aims at minimiz-203 ing variance across a collection of decision trees by averaging their predictions (Breiman, 204 2001). The term 'random' signifies the deliberate characteristic of constructing individ-205 ual trees using different bootstrap samples (sampling observations with replacement) and 206 exploring only a small, randomly selected, subset of the complete input features. This 207 approach effectively decorrelates the individual trees, resulting in a reduction of predic-208 tion variance. Additionally, the construction of each individual tree using only a small 209 subset of input features enables random forest to handle high-dimensional data effectively. 210 The key hyperparameters in a random forest are the number of trees, tree complexity, 211 and the number of randomly selected features used in building the individual trees. Fine-212 tuning these hyperparameters is essential to optimize the performance of the method. 213

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# 2.2.3 Extremely randomized trees

Extremely randomized trees or Extra-trees (ET) operates similarly to RF approach, with the distinction that each tree is constructed using the complete training data, and each split is performed at *random values* using a *random subset* of input features (Geurts et al., 2006). This results in a high degree of independence among the trees and can occasionally yield remarkable results compared to the random forest method.

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# 2.2.4 Adaptive boosting

Adaptive boosting (Adaboost) combines weak learners to create a strong predic-221 tive model (Freund & Schapire, 1997). Weak learners refer to predictive models that per-222 form slightly better than random guesses, and simple decision trees with only a few splits 223 (stumps) are used as weak learners in this study. During each iteration, Adaboost com-224 bines an individual stump by using a weighted sum, where the weight assigned to the 225 current stump is determined based on its overall performance in predicting the target 226 variable. Additionally, the weights associated with the individual training data points 227 are adjusted manually based on their prediction accuracy, giving more attention or weight 228 to points with poor predictions in the next iteration. Adaboost is well known for its abil-229 ity to mitigate overfitting (Rätsch et al., 2001) and has achieved significant success in 230 various prediction challenges (see, for example, Benjamin Bossan (2015) and ZEWEICHU 231 (2019)).232

#### 2.2.5 K-fold cross validation

K-fold cross-validation is the most commonly used model selection technique in machine learning. It involves dividing the training data into K parts or folds, namely  $F_1, \ldots, F_K$ , then a model is trained on K-1 folds, and it is tested on the remaining one. This process is repeated K times and the final performance is the average performance over all the K different testing folds. In this study, K-fold cross-validation is used to prevent overfitting and to select the best possible hyperparameters of each ensemble method. More precisely, if  $f_{\theta}$  is the considered method (random forest, for example) with a hyperparameter  $\theta \in \Theta$ , then the optimal hyperparameter  $\theta^*$  is defined by,

$$\theta^* = \underset{\theta \in \Theta}{\operatorname{argmin}} \frac{1}{K} \sum_{k=1}^{K} \sum_{(x_i, y_i) \in F_k} (f_{\theta}(x_i) - y_i)^2.$$
(1)

In our study,  $\theta$  consists of the depth of the decision trees, the size of random subsets of the ERA5 input features to be considered when building individual trees, and the number of decision trees used in each ensemble learning method. All these keys are tuned using 10-fold cross-validation.

#### 238 2.3 Training

We first train ML models with an extensive set of ERA5 inputs. Subsequently, we refine these inputs to a more manageable subset (see Table 1 below) using importance feature scores, which will be described in Section 3. To reduce the influence of extreme values in the target y and increase its normality, the Box-Cox transformation (Box & Cox, 1964) is performed on the GWMF to obtain the transformed target  $\tilde{y}$ :

$$\tilde{y} = \frac{y^{\lambda} - 1}{\lambda}.$$

In the experiment, the exponent  $\lambda = 0.6$  is chosen based on the performance of models trained on the corresponding transformed target data. The predictions given by ML models are then reverted using the inverse transformation:

$$y = (1 + \lambda \tilde{y})^{1/\lambda}.$$

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Moreover, to predict any GWMFs (absolute, eastward, or westward of HF or WF case) of any given balloon, the ML models are trained using data from the seven other balloons. The models are fine-tuned using a 10-fold cross-validation method to optimize their performances.

## <sup>244</sup> **3 Results**

This section reports the correlations of ML methods in reconstructing various types 245 of observed GWMFs. The numerical study is carried out using sklearn.ensemble mod-246 ule in Python (Pedregosa et al., 2011). In general, the three ML models exhibit similar 247 performances on several balloons, and their variations on any given balloon are insignif-248 icant compared to the variations between different balloons. Additionally, ML models 249 can achieve an encouraging level of correlation larger than 0.7. The average performance 250 over all balloons and data exceeds 0.5, with the westward GWMF showing the worst per-251 formance at a correlation of approximately 0.2. Overall, the performances of ML mod-252 els are sensitive to the choice of balloons and the types of GWs being considered. The 253 numerical results for these cases are provided in the following subsections. 254

<sup>&</sup>lt;sup>1</sup> Solar zenith angle is the only input obtained from the balloons, not from the ERA5. It is a periodic function that provides an estimation of time of the day and the balloon's location.

Name	Notation	Description
Zonal, meridional wind velocity (m/s) & tempera- ture (K)	$\begin{vmatrix} u_j, v_j & \\ temp_j \end{vmatrix}$	with vertical level $j \in \{0, 2, 9, 19\}$ $(km)$ , where 0 is the surface and 19 is the bal- loon's level.
Total precipitation (m)	tp	at center of horizontal grid points.
Mean & standard devia- tion of precipitation (m)	$egin{array}{c} { m tp}_{ m mean} \ \& { m tp}_{ m sd} \end{array}$	over horizontal grid points.
Solar zenith angle (°)	sza <sup>1</sup>	at the location of the balloon.
Log surface pressure (log(hPa))	lnsp	at the surface level.

 Table 1.
 Large-scale input data for training ML models.

## 3.1 Overall performances

Three examples of observed and predicted GWMFs of the HF case are presented 256 in Figure 2 below. Each subplot displays the eastward component of the GWMFs in the 257 positive part and the westward ones in the negative part. It can be observed that the 258 models effectively capture the fluctuations of the observed momentum fluxes, particu-259 larly on balloon 2. However, the models struggle to fully estimate the amplitudes of high-260 peak events, especially for balloons 3 and 7. Overall, the performances of all ML mod-261 els are quite similar; however, there are cases where one outperforms the others. For ex-262 ample, Adaboost appears to do a slightly better job on balloon 2 than the other two mod-263 els in capturing the amplitudes of high-peak events. It is worth noting that balloon 2 264 represents the most satisfying performance of ML methods, balloon 7 is considered the 265 average case, and balloon 3 is the most challenging one to predict. 266

Figure 4 presents boxplots of Pearson's correlation coefficients between predicted 267 and true GWMFs of the HF case. Firstly, choosing the best model is challenging due 268 to the variability in the boxplot positions, which depends on the choices of balloons and 269 GWMF types. For instance, on balloon 2, the correlation boxplot of Adaboost is higher 270 than the other two methods for the absolute and westward cases but lower than Ran-271 dom Forest for the eastward case. However, these differences are generally insignificant 272 compared to the variations observed between different balloons. Secondly, ML models 273 demonstrate strong performance on balloons 2, 6, and 8 across all types of momentum 274 fluxes, and they also excel in predicting the eastward momentum flux of balloon 1. Nev-275 ertheless, balloons 3, 4, 5, and 7 pose greater challenges, with the most difficult being 276 the westward component of GWMF on balloon 3. Finally, the ML models generally out-277 perform the offline gravity wave drag scheme by Lott et al. (2023), except for two cases 278 of balloon 3 (east and westward) and balloon 4. Moreover, Table 4 provides the statis-279 tical significance of the correlations presented in Figure 4. 280

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#### 3.2 Which large-scale inputs are informative for ML models?

The tree-based ensemble ML models employed in this study are not only proficient in predicting GWMFs but also offer valuable insights into the importance of large-scale input information during their training process. Each method exploits the feature im-

F	I AB	$37 \mid \underline{0.43} \mid$	$53 \left  \begin{array}{c} 0.70 \end{array} \right $	23 0.18	$33 \left  \begin{array}{c} 0.37 \end{array} \right $	40 0.50	$\overline{72}$ 0.57	45 0.32	$\overline{66} \mid 0.64 \mid$	y using decorrelated t
	RF E	$0.38 \mid 0.3$	0.60 0.6	$0.21 \mid \underline{0.2}$	$0.35 \mid 0.3$	$0.35 \mid \underline{0.4}$	0.68 0.2	$0.44 \mid \underline{0.4}$	0.66 0.0	each case, by
	AB	0.67	0.65	0.43	0.44	0.35	0.70	0.42	0.68	lution. In
	ET	0.69	0.62	0.49	0.48	0.48	0.65	0.49	0.71	time reso
	RF	0.67	0.67	0.41	0.47	0.39	0.64	0.46	0.71	s in 24h t
	AB	0.58	0.74	0.49	0.47	0.55	0.75	0.48	0.72	GWMF
	ЕT	0.57	0.67	0.48	0.43	0.56	0.74	0.53	0.76	requency
-	RF	0.56	0.70	0.45	0.44	0.51	0.72	0.51	0.74	ed high-f
Duration/ DOF		107/53	103/51	101/33	67/22	79/19	57/10	83/16	77/12	ed and observ
End		28/02/20	23/02/20	28/02/20	02/02/20	23/02/20	01/02/20	28/02/20	22/02/20	tween predict
Start		12/11/19	11/11/19	18/11/19	27/11/19	05/12/19	06/12/19	06/12/19	07/12/19	oefficients be
Alt		20.7	20.2	19.0	18.8	18.9	20.5	20.2	20.2	ation c
Flight		01_STR1	$02\_STR2$	$03_{TTL3}$	$04_{T}TL1$	$05_{-}TTL2$	$06\_STR1$	$07\_STR2$	08_STR2	verage corre

Table 2.	Average correlation coefficients between predicted and observed high-frequency GWMFs in 24h time resolution. In each case, by using decorrelated time
as the degre	ee of freedom (DOF), t-test statistics can provide the significance of each correlation with the convention: <i>italic boldface</i> = $99\%$ , <b>boldface</b> = $95\%$ ,
italic = 90%	%, and normal font = below 90% significant. For any given type of GWMF, the underlined correlations indicate the best performance of ML method on
that target.	



**Figure 2.** Observed and predicted time series of high-frequency east and westward GWMFs of the best, worst and medium cases: balloon 2, 3 and 7, respectively. The x-axis "Day" represents the number of flying days, with 0 corresponding to the moment when the individual balloon was launched.



**Figure 3.** Scatterplots of predictions against observed GWMF corresponding to the time series of Figure 2.

portance (decrement of impurity measure at each split) of its individual decision trees
for determining the overall feature importance, resulting in a ranking of input features
from most to least important. Figure 5 showcases the ranking of the top 5 input features
for all ML methods and GWMF types of the HF case.

Generally, the high-ranking inputs consist of variables that describe precipitation and wind velocity at and below the balloon's level. It is important to note that different models may not rank input features in exactly the same way for a given target (as seen along the rows), due to the variations in the way individual trees are grown. However, the three models concur on the strongly impactful input features; for example, wind velocity at the balloon's level (u19) ranked first in the eastward case (second row) for



**Figure 4.** The boxplots display the correlations between predicted and observed high-frequency GWMFs obtained from 50 runs of ML methods as shown in Table 2. For each balloon, moving from left to right, the three boxplots correspond to the Random forest, Extra-trees, and AdaBoost methods, respectively. The dashed horizontal red lines indicate the performance of the parametrization by Lott et al. (2023).

all models. This suggests that the wind velocity surrounding the balloons is the most
informative large-scale variable for predicting eastward gravity wave momentum fluxes
(GWMFs). Furthermore, the few most significant inputs show a similar preference in both
absolute and eastward GWMFs within the same model, as demonstrated in the columns
of the first and second rows. For instance, standard deviation and average total precipitation (tp\_sd and tp\_mean) are identified as impactful inputs in random forests, while
surface zonal wind velocity (u0) is deemed the most important one in extra trees.

#### 303 4 Discussion

While the results of the machine-learning models are generally encouraging, deficiencies and cases with poor performances were also found. The main motivation for this study being to probe the relationship between the large-scale and the unresolved process, these somewhat negative results are also of interest and can provide useful insights. Possible explanations for the main difficulties encountered are discussed below.

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# 4.1 Why are westward GWMFs more challenging?

Figure 4 displays the performances of the ML models and those of the parameterization used in the IPSL climate model. Balloon 4 constitutes an exception, for which



**Figure 5.** The boxplots show the 5 most important features given by different ML models (by column) on different types of targets (by row). Each boxplot is obtained from the same 50 simulations as displayed in Figure 4.

the parameterization systematically performs better than the ML methods. Leaving bal-312 loon 4 aside, ML approaches unambiguously outperform the parameterization for the ab-313 solute momentum fluxes. For the eastward momentum fluxes, ML approaches generally 314 perform better or are similar to the parameterization. In contrast, both ML approaches 315 and the parameterization have poorer performances for westward MF, and with greater 316 variability for both: for five balloons, ML outperforms clearly the parameterization, whereas 317 for two balloons (including balloon 4) the parameterization clearly outperforms the ML. 318 The present section discusses possible reasons for this difficulty in reproducing the west-319 ward momentum fluxes. 320

Figure 6 displays the Probability Density Function of winds for three balloons as 321 blue curves: balloon 2 has flown in winds that include a majority of westward, strong 322 winds. Like balloon 1, it traveled near  $10^{\circ}$ S in easterly flow for a significant portion of 323 its flight. In contrast, balloons 3 and 7 have flown in weaker winds, with a mild dom-324 inance of westerly winds. Also plotted in figure 6 are conditional PDFs of the zonal winds, 325 conditioned on the intensity of the absolute GWMF. The purpose is to detect if strong 326 values of GWMF were associated to specific wind conditions. For balloon 2, strong GWMF 327 values were found mostly for moderate to strong easterly winds, and this distribution 328 is insensitive to the quantile chosen for the GWMF (90th, 95th or 99th percentile). For 329 balloon 7, the distribution is somewhat sensitive to the quantile chosen. Finally, for bal-330 loon 3, the conditional distribution of zonal wind dramatically changes when it is restricted 331 to the 99th percentile. This detects a particularly intermittent time series, with variabil-332 ity dominated by one extreme event. These findings contribute to explaining the poor 333 performances for balloon 3: the variability of GWMF was dominated there by one (or 334 very few) extreme events, occurring in a specific condition with very weak winds (close 335

# to zero, less than 5 m $s^{-1}$ ). In contrast, the good performances for balloon 2 occur in a case with less intermittency, for which large GWMF are found in strong (easterly) winds.



Figure 6. Conditional densities of zonal wind given different values of high-frequency westward GWMFs. Here, q(0.9), q(0.95) and q(0.99) are the 90%, 95% and 99% quantiles of the absolute value of high-frequency westward GWMFs, respectively.

From table 2, Figure 4 and the trajectories of the balloons (Corcos et al., 2021), it appears that drifting with easterly winds may constitute a favorable factor (balloon 2), but neither a sufficient one (the correlation for westward momentum fluxes for balloon 1, which has a similar trajectory, is moderate, 0.43 at most) nor a necessary one: balloons 6 and 8 generally drift eastward, but good performances are found for the ML reconstruction of the westward MF (0.66 and 0.72 respectively).

Another aspect that influences the performances is the geographical location, and 344 more specifically the latitude of the balloons. Figure 7 displays the PDF of latitude for 345 the eight balloons, distinguishing those for which the ML reconstruction of westward MF 346 is satisfactory (balloons 1, 2, 6 and 8, full lines) from those for which it remains chal-347 lenging (balloons 3, 4, 5 and 7). Here again, one does not isolate a necessary condition, 348 but the balloons for which reconstruction remain challenging are those that remain clos-349 est to the equator. This is consistent with the general expectation that dynamics is more 350 complicated near the Equator, although it is not completely clear why this should mat-351 ter for a small-scale process such as convectively generated gravity waves. It may be that 352 it is not the dynamics itself that is intrinsically more difficult to capture at the Equa-353 tor: it may be the input variables that are poorer, less accurate, very close to the Equa-354 tor. It is known indeed that significant errors, in particular in the wind, are present in 355 the reanalyses very near the Equator (Podglajen et al., 2014; Baker et al., 2014) and the 356 errors are enhanced within a few degrees of the Equator (roughly between  $8^{\circ}S$  and  $8^{\circ}N$ ). 357

# 4.2 Why are some balloons easier to predict than others?

Figure 7 indicates that the predictability of the observed GWMFs is influenced by the balloons' position, specifically, their distances from the equator. Balloons that traveled farther from the equator, primarily south (except for balloon 6, which also explored farther to the north), were found to be easier to predict. This tendency is observed for balloons 1, 2, 6, and 8 which are the well-predicted balloons. In contrast, the challenging balloons spent most of their time flying within a few degrees of the equator, where the atmospheric conditions are not well described by ERA5 data.



**Figure 7.** The trajectories of the balloons during the whole flight (a), and their latitude PDFs (b) and (c). Dashed lines correspond to balloons that pose challenges in prediction.

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## 4.3 Exploring potential reasons for unsatisfactory cases

Several factors could potentially lead to ML models underestimating the observed
 fluxes:

369	•	Part of the relationship between the large-scale flow and a subgrid-scale process
370		such as gravity waves is non-deterministic, or stochastic: for given values of the
371		large-scale fields, a range of different realizations of the subgrid-scale process is
372		possible.
373	•	Input variables from ERA5, which could provide relevant information, may have
374		been omitted.
375	•	The input variables have errors, especially the winds near the Equator (Podglajen
376		et al., 2014). The ML models can not make accurate predictions with an inaccu-
377		rate description of the flow.
378	•	The GWMF are estimated from observations, as a derived quantity; observational
379		error is present in the targeted quantity.
380	•	Given the intermittency of the targeted phenomenon, the sampling by the balloons
381		remains too limited to appropriately train the ML models.

In our study, we mitigated some of the controllable risks mentioned above by first train-382 ing ML models on a large set of ERA5 inputs, then selectively reducing them to a rea-383 sonably small subset, as described in Section 2. This approach ensures that essential ERA5 384 inputs are not inadvertently omitted. Furthermore, fine-tuning the hyperparameters of 385 the models enhances their predictive capacity. Moreover, we observe that all the balloons 386 often flew over many convective processes, and the high-peak events often correspond 387 to deep convective systems, as illustrated for selected cases in Figure 8 below. On Jan-388 uary 12th, 2020, balloon 2 was flying in an area of convection (upper panels (a1) and (a2)), 389 which is likely responsible for the highest peaks in its GWMF timeseries. Interestingly, 390 for balloon 2, almost all events correspond very well with precipitation as described by 391 ERA5 (first column of Figure 9). On the contrary, there is only one big event that hap-392 pened for balloon 3 around January 29th, 2020 (lower left panel (b)). However, the ML 393 models failed to capture it, as it appears to be absent from the ERA5 input variables 394 (not reflected in precipitation nor winds as shown in the second column of Figure 9). This 395 is also true for other challenging balloons, such as the 4th and 5th. Regarding balloon 396 7, the large-scale flows provide partial information for the high-peak events, resulting in 397 partial success in the model's predictions.



**Figure 8.** Brightness temperature from NOAA/NCEP GPM\_MERGIR product (Janowiak, 2017), positions, and the corresponding observed GWMFs at the high-peak events of balloon 2 (top), balloon 3 (lower left) and balloon 7 (lower right).



**Figure 9.** Timeseries of absolute GWMFs and the most informative ERA5 inputs in daily time resolution. The clear correspondence between precipitation and GWMF of balloon 2 can be visually observed in column (a). In contrast, this is not the case at all for balloon 3 as shown in column (b), and it partially presents in column (c) of balloon 7.

# <sup>399</sup> 5 Conclusion and perspectives

# 5.1 Key messages

The relationship between the large-scale atmospheric flow and gravity waves in the lower stratosphere has been investigated using Machine Learning (ML) approaches. This relationship is accounted for in global models through *parameterizations*. ML approaches allow us to revisit these in several ways, notably investigating how much of the subgridscale signal may be estimated *deterministically*, and which are the key variables for that purpose.

Estimates from superpressure balloon measurements were chosen as the target ob-407 servations for gravity wave momentum fluxes (GWMF). The first campaign of the Stra-408 teole 2 project (Haase et al., 2018) consisted of eight balloons flying an average of about 85 days each around the globe in the equatorial band. The quasi-Lagrangian nature of 410 the balloons allows an accurate estimate of gravity wave momentum fluxes (Geller et al., 411 2013), the latter being a key quantity for parameterizations (Alexander et al., 2010). Anal-412 ysis of the GWMF estimated from measurements in this first campaign has highlighted 413 and confirmed convection as the main source of gravity waves in this region, especially 414 for waves with high frequencies (periods shorter than one hour); see Corcos et al. (2021). 415

The description of the large-scale flow environment was provided from the ERA5 reanalysis, along with vertical profiles co-located with each balloon at each time. These variables included wind, pressure, temperature, wind speed, and precipitation. The latter being a noisy and uncertain field, values of total precipitation were retrieved in a 500 × 500 km<sup>2</sup> area around each balloon location, and was generally described by the mean and standard deviation over this area.

The ML methods used focused on tree-based methods: random forests, extremely randomized trees, and adaptive boosting. Other methods were also investigated, as sensitivity experiments, without yielding major improvements. For each method, seven out of eight balloons were used for *training*, and the last balloon was used for *testing*.

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The main results obtained from these investigations are as follows:

427	1.	Based on the information from the large-scale flow described by ERA5, ML meth-
428		ods can reconstruct up to encouraging levels of the observed GWMFs (balloon 2,
429		6, and 8) with correlations larger than 0.7. The performances of ML methods how-
430		ever vary considerably from one balloon to another, with correlations down to 0.4
431		for some other balloons, and even down to 0.2 in one case. The overall average
432		correlation is 0.54.
433	2.	The datasets used for training and testing were more crucial to the performances
434		of the ML approaches than the choice of the ML approaches: for any given bal-
435		loon, the performances of the three methods are mostly similar though not iden-
436		tical, and not always ordered the same way. The tree-based methods proved gen-
437		erally efficient, but there is not an overwhelming preference for one of them. Adap-
438		tive boosting frequently performed a bit better, but all three failed to capture the
439		intensity of the (very intermittent) peaks in GWMF.
440	3.	The most informative explanatory variables are those describing the precipitation
441		and the zonal wind at and below the balloon's level. It is indeed an advantage of
442		tree-based methods to provide information about the usefulness of the different
443		inputs, e.g. through the Gini importance (Hastie et al., 2001). The importance
444		of precipitation is consistent with the convective generation of the waves (Lott $\&$
445		Guez, 2013; Corcos et al., 2021). The importance of winds is consistent with the
446		general understanding of the generation and propagation of waves (Kim et al., 2003);
447		the relevance of wind at the balloon level is reminiscent of previous findings (Plougonven
448		et al., 2017; Amiramjadi et al., 2023).
449	4.	The ML methods were more efficient at reconstructing the part of GWMF asso-
450		ciated with high-frequency waves (periods shorter than an hour) than the whole
451		spectrum. This is consistent with the local character of the explanatory variables
452		provided as inputs: high-frequency waves will be shorter-lived and propagate more
453		vertically. Despite their smaller scales, it is therefore consistent that they are bet-
454		ter reconstructed from local information.
455	5.	To be more precise about the target, different decompositions of the GWMF were
456		used: absolute, eastward and westward GWMF. Interestingly, the performances
457		significantly differed between these. The most difficult to reconstruct was found
458		to be westward GWMF. Reasons for this likely include limitations of the dataset,
459		to be further discussed below.

However, there are still parts where the large-scale flows are not informative enough
in the estimation. There are cases where high peaks are present in the observed target,
which indicates interesting events; however, large-scale flows are missed to describe them.
As a result, the models failed to reconstruct such events in GWMFs (balloon 1 and 3,
for example).

5.2 Perspectives

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Although the ML approaches have performed well, and nearly always better than 466 the parameterization, there are clear limitations to the current investigation, calling for 467 further research. The very strong sensitivity of the performances to the balloon that is 468 left out and then used for testing is a clear indication that we lack data: the results strongly 469 depend on the split of the data for training and testing, the performances are far from 470 convergence. This is consistent with the strong intermittency of the GWMF (Hertzog 471 et al., 2012; Plougonven et al., 2013) and with the illustrative time series of Figure 2: 472 for each balloon, GWMF are dominated by a few events, such that even with 680 days 473 of balloon measurements, only a few handfuls of GWMF peaks are described. This is 474 too little for data-driven methods. This also explains why clear distinctions between the 475 different methods are not found: the ML methods do their best but still lack data to clearly 476 separate a better method for this problem, if there is one. 477

## 478 Ways forward include:

- Obtaining more observations to use as the target, keeping the same framework for the ML. Additional observations would come from the second Strateole 2 campaign (in 2021) and from Loon balloons (Schoeberl et al., 2017; Köhler et al., 2023). The additional Strateole data would enhance the data by less than a factor 2 and is therefore not expected to suffice to make a dramatic change. The Loon data would come with other difficulties as the observations were not made for research purposes and come with their own challenges (Green et al., 2023).
- Additional data could be provided not for the targets, but for the explanatory vari-48F ables. A first step could be including additional input variables from the reanal-487 yses. However, preliminary attempts have not suggested significant gains from the 488 most evident additional culprits. A second step would consist of providing much 489 more detailed and more accurate information about the background flow: this could 490 be obtained from satellite observations, such as the observations of brightness tem-491 peratures from geostationary satellites shown in Figures 8. This would constitute 492 a very interesting new study but in a profoundly new framework and with differ-493 ent aims: to fully use the information available from satellites would a priori require providing maps (or images, or 2D fields) as input variables (more akin to 495 Matsuoka et al. (2020), although their inputs were from models, not observations) 496 The ML used would need to be reassessed (Matsuoka et al. (2020) used neural net-497 works, for instance). Such a study would be of great interest because the perfor-498 mance of the ML methods would much less be tainted by the uncertainty (or er-499 rors) present in the inputs that serve to describe the background. Additionally, 500 much more detailed information would be provided about the background flow, 501 allowing the ML methods to tap into a greater reservoir of potentially relevant in-502 formation, and hence providing more precise answers regarding the relationship 503 of the large- scale flow to the gravity wave signal. However, if the outcome of such 504 an exercise would be of interest fundamentally, it would be more removed from 505 the framework in which current parameterizations operate. 506
- A third way forward consists of applying similar investigations on datasets where 507 more data is available, albeit at the cost of more uncertainty on the realism of the 508 data. High-resolution models such as global convection permitting simulations (Stephan 509 et al., 2019) provide a wealth of information on the resolved gravity wave field, 510 and many studies have repeatedly highlighted the ability of models to simulate 511 efficiently many features of the observed gravity wave field (Plougonven & Teit-512 elbaum, 2003; Wu & Eckermann, 2008; Preusse et al., 2014; Stephan et al., 2019). 513 Model output from global simulations would provide amounts of data for which 514 the sampling limitations of the Strateole balloons would not be present. The down-515 side is the limitations of model data, relative to observations, and the need for strate-516 gies to validate which aspects of the simulations are realistic. 517

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# 525 6 Open research

Balloon data used in this study are presented in Haase et al. (2018) of the STRA-TEOLE 2 mission and can be extracted from the following website: https://webstr2 <sup>528</sup> .ipsl.polytechnique.fr. The ERA5 input variables are described in Hersbach et al.

- (2020) and can be obtained from the COPERNICUS access hub using the following web-
- site: https://scihub.copernicus.eu/. The machine learning algorithms implemented
- in our analysis are available in the scikit-learn python library (Pedregosa et al., 2011)
- and can be downloaded from its website: https://scikit-learn.org/stable/install
- .html. Finally, the source codes for implementing machine learning methods in our anal-
- ysis are made available at the following GitHub repository: https://github.com/hassothea/
   Reconstruction\_of\_GWMF\_using\_ML\_ERA5.

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